## The Global Stability of the Minkowski Spacetime Solution to the Einstein-Nonlinear Electromagnetic System in Wave Coordinates

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## The "Electromagnetic Divergence Problem"

Intro

- Electric field surrounding an electron:  $\mathbf{E}_{(Maxwell)} = -rac{\hat{\mathbf{r}}}{|\mathbf{r}|^2}$
- Electrostatic energy =  $\int_{\mathbb{R}^3} |\mathbf{E}_{(Maxwell)}|^2 dVol$ =  $4\pi \int_{r=0}^{\infty} \frac{1}{r^2} dr = \infty$
- Lorentz force: F<sub>(Lorentz)</sub> = qE<sub>(Maxwell)</sub> = Undefined (at the electron's location)
- Analogous difficulties appear in QED

M. Born: "The attempts to combine Maxwell's equations with the quantum theory (...) have not succeeded. One can see that the failure does not lie on the side of the quantum theory, but on the side of the field equations, which do not account for the existence of a radius of the electron (or its finite energy=mass)."

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## Towards a possible resolution

- Idea (Born, Infeld, Bialynicki-Birula, Carley, Kiessling, Tahvildar-Zadeh, · · · ): nonlinear field equations
- Example: the Born-Infeld model ( finite energy electrons)
- Lorentz force remains undefined; Kiessling → new law of motion for point charges
- $\bullet$  All physically reasonable models  $\sim$  Maxwell model in the weak-field limit
- Minkowski spacetime (in absence of point charges). This is what my work confirms.

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- All physically reasonable models ~ Maxwell model in the weak-field limit

No physically reasonable model should destroy the stability of Minkowski spacetime (in absence of point charges). **This is what my work confirms**.

# The 1 $\pm$ 3 dimensional Einstein-nonlinear electromagnetic equations

$$egin{aligned} \mathsf{Ric}_{\mu
u} - rac{1}{2} g_{\mu
u} R = T_{\mu
u}, \ (d\mathcal{F})_{\lambda\mu
u} = 0, \ (d\mathcal{M})_{\lambda\mu
u} = 0. \end{aligned}$$

- $g_{\mu\nu}$  is the spacetime metric;  $\mathrm{Ric}_{\mu\nu} = \mathrm{Ric}_{\mu\nu}(g,\partial g,\partial^2 g)$  is the Ricci tensor;  $R \stackrel{\text{def}}{=} (g^{-1})^{\kappa\lambda} \mathrm{Ric}_{\kappa\lambda}$  is the scalar curvature
- $T_{\mu\nu}$  is the energy-momentum tensor
- $\mathcal{F}_{\mu\nu}$  is the Faraday tensor
- $\mathcal{M}_{\mu\nu}$  is the Maxwell tensor

## Previous mathematical results in 1 + 3 dimensions

- Fritz John: blow-up for nonlinear wave equations in three spatial dimensions (1981)
- Klainerman: global existence for nonlinear wave equations and the null condition (1984)
- Friedrich: "stability of the Minkowski spacetime solution" to the Einstein-vacuum equations for restricted data using the conformal method (1986)
- Christodoulou-Klainerman: " "using an invariant framework (1993)
- Nina Zipser: " " to the Einstein-Maxwell system (2000)
- Lindblad-Rodnianski: " " to the Einstein-scalar field system in wave coordinates (2005, 2010)
- Julian Loizelet: " " to the Einstein-scalar field-Maxwell system in Lorenz-gauge + wave coordinates (2008)
- JS: " to the Einstein-nonlinear electromagnetic system in wave coordinates (2010)



## The initial data on $\Sigma_0 = \mathbb{R}^3$

 $\underline{\mathring{g}}=$  Riemannian metric

 $\overline{\mathring{K}} = \mathsf{symmetric} \; \mathsf{type} \; \left( \begin{smallmatrix} 0 \\ 2 \end{smallmatrix} \right) \; \mathsf{tensor} \; .$ 

 $\mathring{\mathfrak{D}} = \text{electric displacement}$ 

 $\mathring{\mathfrak{B}} = \mathsf{magnetic} \; \mathsf{induction}$ 

Constraint equations (j = 1, 2, 3)

$$\begin{split} & \underline{\mathring{R}} - \mathring{K}_{ab}\mathring{K}^{ab} + \left[ (\underline{\mathring{g}}^{-1})^{ab}\mathring{K}_{ab} \right]^2 = 2T(\hat{N}, \hat{N})|_{\Sigma_0}, \\ & (\underline{\mathring{g}}^{-1})^{ab}\underline{\mathring{\mathcal{D}}}_{a}\mathring{K}_{bj} - (\underline{\mathring{g}}^{-1})^{ab}\underline{\mathring{\mathcal{D}}}_{j}\mathring{K}_{ab} = T(\hat{N}, \frac{\partial}{\partial x^{j}})|_{\Sigma_0}, \\ & (\underline{\mathring{g}}^{-1})^{ab}\underline{\mathring{\mathcal{D}}}_{a}\mathring{\mathfrak{D}}_{b} = 0, \\ & (\underline{\mathring{g}}^{-1})^{ab}\underline{\mathring{\mathcal{D}}}_{a}\mathring{\mathfrak{B}}_{b} = 0 \end{split}$$



## Positive mass theorem of Schoen-Yau, Witten

Decay assumptions on the data ( $\kappa > 0$ , M > 0)

$$\bullet \ \underline{\mathring{g}}_{jk} = \delta_{jk} + \underline{\mathring{h}}_{jk}^{(0)} + \underline{\mathring{h}}_{jk}^{(1)},$$

$$\bullet \; \underline{\mathring{h}}_{jk}^{(0)} = \chi(r) \frac{2M}{r} \delta_{jk},$$

$$\bullet \ \underline{\mathring{h}}_{ik}^{(1)} = o(r^{-1-\kappa}),$$

$$\frac{n_{jk}}{N} = O(1)$$

• 
$$\mathring{K}_{ik} = o(r^{-2-\kappa}),$$

• 
$$\mathring{\mathfrak{D}}_i = o(r^{-2-\kappa}),$$

• 
$$\mathring{\mathfrak{B}}_i = o(r^{-2-\kappa}),$$

as 
$$r \to \infty$$
.

as 
$$r \to \infty$$
.

as 
$$r \to \infty$$
.

as 
$$I \to \infty$$
,

as 
$$r \to \infty$$

## Splitting the spacetime metric

#### Minkowski metric

$$ullet g_{\mu
u} = egin{pmatrix} \widehat{m}_{\mu
u} & +h_{\mu
u} \end{pmatrix}$$

$$\bullet \ h_{\mu\nu} = h_{\mu\nu}^{(0)} + h_{\mu\nu}^{(1)}$$

Schwarzschild tail remainder piece

• 
$$h_{\mu\nu}^{(0)} = \chi(\frac{r}{t})\chi(r)\frac{2M}{r}\delta_{\mu\nu}$$

$$\chi \in C^{\infty}$$
,  $\chi \equiv 1$  for  $z \geq 3/4$ ,  $\chi \equiv 0$  for  $z \leq 1/2$ .

Note:  $\chi'(\frac{r}{t})$  is supported in the interior region  $\{[1/2 \le r/t \le 3/4]\}$ 

## Assumptions on the electromagnetic Lagrangian

$$\bullet \ \ \ \, \zeta_{(1)}\stackrel{\mathsf{def}}{=} \frac{1}{2} (g^{-1})^{\kappa\mu} (g^{-1})^{\lambda\nu} \mathcal{F}_{\kappa\lambda} \mathcal{F}_{\mu\nu} \simeq \text{``} |\mathbf{B}|^2 - |\mathbf{E}|^2, \text{''}$$

• 
$$abla_{(2)} \stackrel{\text{def}}{=} \frac{1}{4} (g^{-1})^{\kappa\mu} (g^{-1})^{\lambda\nu} \mathcal{F}_{\kappa\lambda} {}^{\star} \mathcal{F}_{\mu\nu} = \frac{1}{8} \epsilon^{\#\kappa\lambda\mu\nu} \mathcal{F}_{\kappa\lambda} \mathcal{F}_{\mu\nu} \simeq$$
"**E** · **B**"

$$\left| \mathcal{T}_{\kappa\lambda} X^{\kappa} Y^{\lambda} \ge 0 \right|$$
 (Dominant energy condition) (A2)

#### whenever

- X, Y are both timelike (i.e.,  $g_{\kappa\lambda}X^{\kappa}X^{\lambda} < 0$ ,  $g_{\kappa\lambda}Y^{\kappa}Y^{\lambda} < 0$ )
- X. Y are future-directed



## Example: the Born-Infeld Lagrangian

$$^{\star} \mathcal{L}_{(BI)} = 1 - \sqrt{1 + 4_{(1)} - 4_{(2)}^{2}} 
= -\frac{1}{2} 4_{(1)} + O(|(4_{(1)}, 4_{(2)})|^{2})$$

## The Maxwell tensor and the energy-momentum tensor

$$\bullet \mathcal{M}_{\mu\nu} = 2 \frac{\partial^* \mathcal{L}}{\partial \zeta_{(1)}} {}^* \mathcal{F}_{\mu\nu} - \frac{\partial^* \mathcal{L}}{\partial \zeta_{(2)}} \mathcal{F}_{\mu\nu} 
= \underbrace{{}^* \mathcal{F}_{\mu\nu}}_{} + O(|h||\mathcal{F}|) + O(|\mathcal{F}|^3; h),$$

Minkowskian Hodge dual

• 
$$T_{\mu\nu} = -2 \frac{\partial^* \mathcal{L}}{\partial \dot{\gamma}_{(1)}} (g^{-1})^{\kappa\lambda} \mathcal{F}_{\mu\kappa} \mathcal{F}_{\nu\lambda} - \dot{\gamma}_{(2)} \frac{\partial^* \mathcal{L}}{\partial \dot{\gamma}_{(2)}} g_{\mu\nu} + g_{\mu\nu}^* \mathcal{L}$$

$$= \underbrace{(m^{-1})^{\kappa\lambda} \mathcal{F}_{\mu\kappa} \mathcal{F}_{\nu\lambda} - \frac{1}{4} m_{\mu\nu} (m^{-1})^{\kappa\eta} (m^{-1})^{\lambda\zeta} \mathcal{F}_{\kappa\lambda} \mathcal{F}_{\eta\zeta}}_{\text{Minkowskian/linear energy-momentum tensor}} + O(|h||\mathcal{F}|^2) + O(|\mathcal{F}|^3; h)$$

## The wave coordinate gauge

Wave coordinates (à la Choquet-Bruhat & Lindblad-Rodnianski):

$$\bullet (g^{-1})^{\kappa\lambda}\Gamma^{\mu}_{\kappa\lambda}=0$$

$$\implies \mathsf{Ric}_{\mu 
u} = -rac{1}{2}\widetilde{\Box}_g g_{\mu 
u} + \mathcal{N}_{\mu 
u}(g, 
abla g)$$

- ullet  $\widetilde{\Box}_g = (g^{-1})^{\kappa\lambda} \nabla_{\kappa} \nabla_{\lambda}$
- $\bullet$   $\nabla$  = Minkowski connection

## The reduced equations

$$egin{aligned} \widetilde{\Box}_{g}h_{\mu
u}^{(1)} &= \mathfrak{H}_{\mu
u} - \widetilde{\Box}_{g}h_{\mu
u}^{(0)}, \ 
abla_{\lambda}\mathcal{F}_{\mu
u} + 
abla_{\mu}\mathcal{F}_{
u\lambda} + 
abla_{
u}\mathcal{F}_{\lambda\mu} &= 0, \ N^{\#\mu
u\kappa\lambda}
abla_{\mu}\mathcal{F}_{\kappa\lambda} &= \mathfrak{F}^{
u}. \end{aligned}$$

$$\begin{split} \bullet \, \mathfrak{H}_{\mu\nu} &= \mathscr{P}(\nabla_{\mu}h, \nabla_{\nu}h) + \mathscr{Q}_{\mu\nu}^{(1;h)}(\nabla h, \nabla h) + \mathscr{Q}_{\mu\nu}^{(2;h)}(\mathcal{F}, \mathcal{F}) + \mathfrak{H}_{\mu\nu}^{\triangle}, \\ \bullet \, \mathfrak{F}^{\nu} &= \mathscr{Q}_{(2;\mathcal{F})}^{\nu}(\nabla h, \mathcal{F}) + \mathfrak{F}_{\triangle}^{\nu}, \\ \bullet \, N^{\#\mu\nu\kappa\lambda} &= -\frac{1}{2} \frac{\partial^{2}\mathscr{L}}{\partial \mathcal{F}_{\mu\nu}\partial \mathcal{F}_{\kappa\lambda}} + \frac{1}{2} \frac{\partial^{*}\mathscr{L}}{\partial \mathcal{F}_{(2)}} \epsilon^{\#\mu\nu\kappa\lambda} \\ &= \frac{1}{2} \Big\{ (m^{-1})^{\mu\kappa} (m^{-1})^{\nu\lambda} - (m^{-1})^{\mu\lambda} (m^{-1})^{\nu\kappa} \\ &\qquad \qquad - 2h^{\mu\kappa} (m^{-1})^{\nu\lambda} - 2(m^{-1})^{\mu\kappa} h^{\nu\lambda} \Big\} + N_{\triangle}^{\#\mu\nu\kappa\lambda}. \end{split}$$

## The energy of the reduced solution

$$\bullet \,\, \mathcal{E}^2_{\ell;\gamma;\mu}(t) \stackrel{\text{def}}{=} \sup_{0 \leq \tau \leq t} \sum_{|I| < \ell} \int_{\Sigma_{\mathcal{I}}} \Big\{ |\nabla \nabla^I_{\mathcal{Z}} h^{(1)}|^2 + |\mathcal{L}^I_{\mathcal{Z}} \mathcal{F}|^2 \Big\} w(q) \, d^3x,$$

$$q\stackrel{\text{def}}{=} |x|-t,$$

• 
$$w = w(q) = \begin{cases} 1 + (1 + |q|)^{1+2\gamma}, & \text{if } q > 0, \\ 1 + (1 + |q|)^{-2\mu}, & \text{if } q < 0, \end{cases}$$

$$0 < \gamma < 1/2, \, 0 < \mu < 1/2,$$

• 
$$\mathcal{Z} \stackrel{\text{def}}{=} \left\{ \frac{\partial}{\partial x^{\mu}}, x_{\mu} \frac{\partial}{\partial x^{\nu}} - x_{\nu} \frac{\partial}{\partial x^{\mu}}, x^{\kappa} \frac{\partial}{\partial x^{\kappa}} \right\}_{0 \leq \mu < \nu \leq 3}$$

11 Minkowskian conformal Killing fields



## Continuation principle

#### Lemma

If the solution blows up at time  $T_{max} > 0$ , then one of the following blow-up scenarios must occur:

- The hyperbolicity of the equations breaks down as t ↑ T<sub>max</sub>
- $\lim_{t \uparrow T_{max}} \mathcal{E}_{\ell;\gamma;\mu}(t) = \infty$

Moral conclusion: You can show global existence by proving that the energy  $\mathcal{E}_{\ell;\gamma;\mu}(t)$  never blows up.

## Main results

#### Theorem

Let  $\ell \geq 8$  and assume the wave coordinate condition. There exists a constant  $\varepsilon_{\ell} > 0$  such that if  $\mathcal{E}_{\ell;\gamma;\mu}(0) + M \leq \varepsilon \leq \varepsilon_{\ell}$ , then the reduced solution exists for  $(t,x) \in (-\infty,\infty) \times \mathbb{R}^3$  and the resulting spacetime is geodesically complete. Furthermore, there exist constants  $c_{\ell} > 0$ ,  $\widetilde{c}_{\ell} > 0$  such that

$$\mathcal{E}_{\ell;\gamma;\mu}(t) \leq c_\ell \varepsilon (1+|t|)^{\widetilde{c}_\ell \varepsilon}$$

for  $t \in (-\infty, \infty)$ .

## The bootstrap argument

Goal: assume that there exist constants  $0 < \varepsilon$  and  $0 < \delta < 1/4$  such that for  $t \in [0, T)$ 

• 
$$\mathcal{E}_{\ell;\gamma;\mu}(t) \leq \varepsilon (1+t)^{\delta}$$
,  $0 \leq k \leq \ell$ ,

and prove that for  $0 \le k \le \ell$  and  $t \in [0, T)$ :

$$\begin{split} \mathcal{E}_{k;\gamma;\mu}^2(t) &\lesssim \mathcal{E}_{\ell;\gamma;\mu}^2(0) + M^2 + \varepsilon^3 + \varepsilon \int_0^t (1+\tau)^{-1} \mathcal{E}_{k;\gamma;\mu}^2(\tau) \, d\tau \\ &+ \varepsilon \int_0^t (1+\tau)^{-1+C\varepsilon} \mathcal{E}_{k-1;\gamma;\mu}^2(\tau) \, d\tau \end{split}$$

Conclusion (Gronwall's lemma):

$$\mathcal{E}^2_{\ell:\gamma:u}(t) \leq c_\ell \{\mathcal{E}^2_{\ell:\gamma:u}(0) + M^2 + \varepsilon^3\} (1+t)^{\widetilde{c}_\ell \varepsilon}$$



## Fundamental difficulty

The behavior of  $\mathcal{E}^2_{k;\gamma;\mu}(t)$  is not directly accessible. We only know how to study  $\mathcal{E}^2_{k;\gamma;\mu}(t)$  through the use of auxiliary quantities, e.g. energy currents.

## The EOV and the canonical stress

The equations of variation (EOV)

$$abla_{\lambda}\dot{\mathcal{F}}_{\mu
u} + 
abla_{\mu}\dot{\mathcal{F}}_{\nu\lambda} + 
abla_{
u}\dot{\mathcal{F}}_{\lambda\mu} = 
abla_{\lambda\mu\nu}^{0 ext{ for us}},$$
 $abla^{\mu\mu\nu\kappa\lambda}[h,\mathcal{F}]\nabla_{\mu}\dot{\mathcal{F}}_{\kappa\lambda} = \dot{\mathfrak{F}}^{\nu}$ 

- To estimate  $\mathcal{F}$ , we can use  $(g^{-1})^{\kappa\lambda}\mathscr{D}_{\kappa}T_{\mu\lambda}=0$ .
- Answer: the canonical stress tensor (Christodoulou)

$$\dot{Q}^{\mu}_{\ \nu} \stackrel{\mathrm{def}}{=} N^{\#\mu\zeta\kappa\lambda} \dot{\mathcal{F}}_{\kappa\lambda} \dot{\mathcal{F}}_{\nu\zeta} - \frac{1}{4} \delta^{\mu}_{\ \nu} N^{\#\zeta\eta\kappa\lambda} \dot{\mathcal{F}}_{\zeta\eta} \dot{\mathcal{F}}_{\kappa\lambda}$$

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u} + 
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abla_{
u}\dot{\mathcal{F}}_{\lambda\mu} = \widehat{\dot{\mathfrak{F}}}_{\lambda\mu\nu} \ , \ N^{\#\mu
u\kappa\lambda}[h,\mathcal{F}]
abla_{\mu}\dot{\mathcal{F}}_{\kappa\lambda} = \dot{\mathfrak{F}}^{
u}$$

- To estimate  $\mathcal{F}$ , we can use  $(g^{-1})^{\kappa\lambda}\mathscr{D}_{\kappa}T_{\mu\lambda}=0$ .
- But how do we obtain useful differential identities for  $\dot{\mathcal{F}}$ ?
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$$\boxed{ \dot{Q}^{\mu}_{\ \nu} \stackrel{\text{def}}{=} \textit{\textbf{N}}^{\#\mu\zeta\kappa\lambda} \dot{\mathcal{F}}_{\kappa\lambda} \dot{\mathcal{F}}_{\nu\zeta} - \frac{1}{4} \delta^{\mu}_{\nu} \textit{\textbf{N}}^{\#\zeta\eta\kappa\lambda} \dot{\mathcal{F}}_{\zeta\eta} \dot{\mathcal{F}}_{\kappa\lambda} }$$

## Expansion of the canonical stress

linear energy-momentum tensor in flat spacetime

$$\dot{Q}^{\mu}_{\ 
u} = \qquad \qquad \dot{\dot{\mathcal{F}}^{\mu\zeta}\dot{\mathcal{F}}_{\nu\zeta} - rac{1}{4}\delta^{\mu}_{
u}\dot{\mathcal{F}}_{\zeta\eta}\dot{\mathcal{F}}^{\zeta\eta}}$$

corrections to linear theory arising from h

$$-h^{\mu\kappa}\dot{\mathcal{F}}_{\kappa\zeta}\dot{\mathcal{F}}_{\nu}^{\ \zeta} - h^{\kappa\lambda}\dot{\mathcal{F}}_{\kappa}^{\mu}\dot{\mathcal{F}}_{\nu\lambda} + \frac{1}{2}\delta_{\nu}^{\mu}h^{\kappa\lambda}\dot{\mathcal{F}}_{\kappa\eta}\dot{\mathcal{F}}_{\lambda}^{\ \eta} + \underbrace{N_{\triangle}^{\#\mu\zeta\kappa\lambda}\dot{\mathcal{F}}_{\kappa\lambda}\dot{\mathcal{F}}_{\nu\zeta} - \frac{1}{4}\delta_{\nu}^{\mu}N_{\triangle}^{\#\zeta\eta\kappa\lambda}\dot{\mathcal{F}}_{\zeta\eta}\dot{\mathcal{F}}_{\kappa\lambda}}_{\text{quartic error terms}}.$$

## An expression for $abla_{\mu}\dot{Q}^{\mu}_{\ u}$

#### Lemma

If  $\dot{\mathcal{F}}_{\mu\nu}$  verifies the equations of variation with inhomogeneous terms  $\dot{\mathfrak{F}}_{\nu\kappa\lambda}$  and  $\dot{\mathfrak{F}}^{\eta}$ , then

$$\begin{split} \nabla_{\mu} \dot{Q}^{\mu}_{\ \nu} &= -\frac{1}{2} N^{\#\zeta\eta\kappa\lambda} \dot{\mathcal{F}}_{\zeta\eta} \dot{\mathfrak{F}}_{\nu\kappa\lambda} + \dot{\mathcal{F}}_{\nu\eta} \dot{\mathfrak{F}}^{\eta} \\ &- (\nabla_{\mu} h^{\mu\kappa}) \dot{\mathcal{F}}_{\kappa\zeta} \dot{\mathcal{F}}_{\nu}^{\ \zeta} - (\nabla_{\mu} h^{\kappa\lambda}) \dot{\mathcal{F}}^{\mu}_{\kappa} \dot{\mathcal{F}}_{\nu\lambda} + \frac{1}{2} (\nabla_{\nu} h^{\kappa\lambda}) \dot{\mathcal{F}}_{\kappa\eta} \dot{\mathcal{F}}^{\eta}_{\lambda} \\ &+ (\nabla_{\mu} N^{\#\mu\zeta\kappa\lambda}_{\triangle}) \dot{\mathcal{F}}_{\kappa\lambda} \dot{\mathcal{F}}_{\nu\zeta} - \frac{1}{4} (\nabla_{\nu} N^{\#\zeta\eta\kappa\lambda}_{\triangle}) \dot{\mathcal{F}}_{\zeta\eta} \dot{\mathcal{F}}_{\kappa\lambda} \end{split}$$

## A current for $\dot{\mathcal{F}}$

Set  $X^{
u}=w(q)\delta_0^{
u}$  and define

$$\dot{J}^{\mu}_{(h,\mathcal{F})}[\dot{\mathcal{F}}] \stackrel{\mathsf{def}}{=} -\dot{Q}^{\mu}_{\ 
u} X^{
u} = -w(q)\dot{Q}^{\mu}_{\ 0}$$

$$w(q) \stackrel{\text{def}}{=} \left\{ egin{array}{ll} 1 + (1 + |q|)^{1 + 2\gamma}, & ext{if } q > 0, \ 1 + (1 + |q|)^{-2\mu}, & ext{if } q < 0, \end{array} 
ight.$$

$$0 < \gamma < 1/2, \quad 0 < \mu < 1/2.$$

#### Lemma

$$\dot{J}^0_{(h,\mathcal{F})}[\dot{\mathcal{F}}] \approx |\dot{\mathcal{F}}|^2 w(q).$$

## The divergence theorem

#### Lemma

Let  $\dot{\mathcal{F}}_{\mu\nu}$  be a solution to the electromagnetic equations of variation with inhomogeneous terms  $\dot{\mathfrak{F}}^{\nu}$ . Assume that " $h\lesssim \varepsilon$ " is suitably small. Then

$$\begin{split} \int_{\Sigma_t} |\dot{\mathcal{F}}|^2 w(q) \, d^3 x \, + \, \int_0^t \int_{\Sigma_\tau} \left( |\dot{\alpha}|^2 + \dot{\rho}^2 + \dot{\sigma}^2 \right) w'(q) \, d^3 x \, d\tau \\ \lesssim \int_{\Sigma_0} |\dot{\mathcal{F}}|^2 w(q) \, d^3 x \, + \, \varepsilon \int_0^t \int_{\Sigma_\tau} \frac{|\dot{\mathcal{F}}|^2}{1 + \tau} w(q) \, d^3 x \, d\tau \\ + \, \int_0^t \int_{\Sigma_\tau} |\dot{\mathcal{F}}_{0\kappa} \dot{\mathfrak{F}}^\kappa | w(q) \, d^3 x \, d\tau \end{split}$$

# Weighted Klainerman-Sobolev inequality (with $q \stackrel{\text{def}}{=} |x| - t$ )

#### Lemma

$$(1+t+|x|)[(1+|q|)w(q)]^{1/2}|\phi(t,x)| \lesssim \sum_{|I| \leq 2} \|w^{1/2}\nabla_{\mathcal{Z}}^{I}\phi(t,\cdot)\|_{L^{2}},$$

$$(1+t+|x|)[(1+|q|)w(q)]^{1/2}|\mathcal{F}(t,x)| \lesssim \sum_{|I| \leq 2} \|w^{1/2}\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}(t,\cdot)\|_{L^{2}},$$

$$(1+t+|x|)^{2}[(1+|q|)w(q)]^{1/2}|\overline{\nabla}\phi(t,x)| \lesssim \sum_{|I| \leq 3} \|w^{1/2}\nabla_{\mathcal{Z}}^{I}\phi(t,\cdot)\|_{L^{2}},$$

Not sufficient to close the estimates; we need to upgrade these inequalities using the **special null structure** of the equations.

## Weighted Klainerman-Sobolev inequality (with $q \stackrel{\text{def}}{=} |x| - t$ )

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$$(1+t+|x|)[(1+|q|)w(q)]^{1/2}|\mathcal{F}(t,x)| \lesssim \sum_{|I|\leq 2} \|w^{1/2}\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}(t,\cdot)\|_{L^{2}},$$

$$(1+t+|x|)^{2}[(1+|q|)w(q)]^{1/2}|\overline{\nabla}\phi(t,x)| \lesssim \sum_{|I|\leq 3} \|w^{1/2}\nabla_{\mathcal{Z}}^{I}\phi(t,\cdot)\|_{L^{2}},$$

Not sufficient to close the estimates; we need to upgrade these inequalities using the **special null structure** of the equations.

## Heuristics

Moral reason for stability: the "worst possible" quadratic terms are absent from the equations

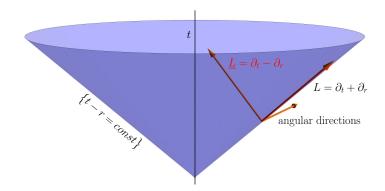
weak null condition

## Important surfaces in Minkowski space

$$t \in \mathbb{R}, \ x \in \mathbb{R}^3$$

- $C_s^- \stackrel{\text{def}}{=} \{(\tau, y) \mid |y| + \tau = s\}$  are the ingoing null cones
- $C_q^+\stackrel{\text{def}}{=} \{(\tau,y) \mid |y|-\tau=q\}$  are the outgoing null cones
- $\Sigma_t \stackrel{\text{def}}{=} \{(\tau, y) \mid \tau = t\}$  are the constant time slices
- $S_{r,t} \stackrel{\text{def}}{=} \{(\tau, y) \mid \tau = t, |y| = r\}$  are the Euclidean spheres

### The Minkowskian null frame



## Minkowskian null frame and null coordinates

Null frame:  $\mathcal{N} \stackrel{\text{def}}{=} \{\underline{L}, L, e_1, e_2\}, \ \mathcal{T} \stackrel{\text{def}}{=} \{L, e_1, e_2\}, \ \mathcal{L} \stackrel{\text{def}}{=} \{L\}$ 

- $\underline{L} \stackrel{\text{def}}{=} \partial_t \partial_r$  is tangent to the ingoing cones
- $L \stackrel{\text{def}}{=} \partial_t + \partial_r$  is tangent to the outgoing cones
- $e_1, e_2$  are orthonormal, & tangent to the spheres

Null coordinates (useful for expressing decay rates)

- q = r t (constant on outgoing cones)
- s = r + t (constant on ingoing cones)

## Null decomposition of $\mathcal{F}$

With 
$$m_{\mu\nu}\stackrel{\text{def}}{=} m_{\mu\nu} + \frac{1}{2} (L_{\mu}\underline{L}_{\nu} + \underline{L}_{\mu}L_{\nu}), \ \psi_{\mu\nu}\stackrel{\text{def}}{=} \frac{1}{2} [\mu\nu\kappa\lambda]\underline{L}^{\kappa}L^{\lambda},$$

- $m_{\mu}^{\nu}$  projects m-orthogonally onto the  $S_{r,t}$
- ullet  $\pi_{\mu}^{\ 
  u}$  projects m-orthogonally onto the  $C_q^+$
- $\overline{\nabla}_{\mu} \stackrel{\text{def}}{=} \overline{\pi}_{\mu}^{\ \kappa} \nabla_{\kappa}$  are the good derivatives

## Upgraded decay estimates via wave coordinates

Null decompose the wave coordinate condition  $\nabla_{\mu}(\sqrt{|\det g|}(g^{-1})^{\mu\nu})=0$  to obtain (Lindblad-Rodnianski)

$$\boxed{|\nabla h|_{\mathcal{LT}} \lesssim |\overline{\nabla} h| + |h||\nabla h|}$$

Conclusion:  $|h|_{CT}$  decays better than "expected"

## Upgraded decay estimates for $\mathcal{F}$

With  $\underline{\alpha}_{\mu} \stackrel{\text{def}}{=} p n_{\mu}^{\nu} \mathcal{F}_{\nu\lambda} \underline{L}^{\lambda}$ ,  $\Lambda \stackrel{\text{def}}{=} L + \frac{1}{4} h_{LL} \underline{L}$ ,

Null decomposing the electromagnetic equations  $\Longrightarrow$ 

$$\begin{split} r^{-1}|\nabla_{\Lambda}(r\underline{\alpha})| &\lesssim r^{-1}|h|_{\mathcal{LL}}|\underline{\alpha}| \ + \ \sum_{|I| \leq 1} r^{-1} \big(|\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}|_{\mathcal{LN}} + |\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}|_{\mathcal{TT}}\big) \\ &+ \sum_{|I_{1}| + |I_{2}| \leq 1} r^{-1}|\nabla_{\mathcal{Z}}^{I_{1}}h||\mathcal{L}_{\mathcal{Z}}^{I_{2}}\mathcal{F}| \\ &+ \sum_{|I| \leq 1} (1+|q|)^{-1}|h| \big(|\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}|_{\mathcal{LN}} + |\mathcal{L}_{\mathcal{Z}}^{I}\mathcal{F}|_{\mathcal{TT}}\big) \\ &+ \text{ cubic error terms.} \end{split}$$

- Conclusion: r
   <u>α</u> satisfies an "ODE" with favorable sources
- Integration yields  $|\alpha| \le \epsilon (1+t)^{-1}$
- Klainerman-Sobolev yields only  $|\alpha| \lesssim \epsilon (1+t)^{-1+\delta}$

## Post-upgraded-decay-estimate integral inequalities

$$\begin{split} \sum_{|I| \leq k} \int_0^t \int_{\Sigma_{\tau}} \Big| (\mathcal{L}_{\mathcal{Z}}^I \mathcal{F}_{0\nu}) \Big\{ N^{\#\mu\nu\kappa\lambda} \nabla_{\mu} \mathcal{L}_{\mathcal{Z}}^I \mathcal{F}_{\kappa\lambda} - \hat{\mathcal{L}}_{\mathcal{Z}}^I \left( N^{\#\mu\nu\kappa\lambda} \nabla_{\mu} \mathcal{F}_{\kappa\lambda} \right) \Big\} \Big| w(q) \, d^3x \, d\tau \\ &\lesssim \varepsilon \sum_{|I| \leq k} \int_0^t \int_{\Sigma_{\tau}} (1+\tau)^{-1} |\mathcal{L}_{\mathcal{Z}}^I \mathcal{F}|^2 w(q) \, d^3x \, d\tau \\ &+ \varepsilon \sum_{|I| \leq k} \int_0^t \int_{\Sigma_{\tau}} (1+\tau)^{-1} |\nabla [\nabla_{\mathcal{Z}}^I h^{(1)}]|^2 w(q) \, d^3x \, d\tau \\ &+ \varepsilon \sum_{|I| \leq k} \int_0^t \int_{\Sigma_{\tau}} \Big\{ |\overline{\nabla} [\nabla_{\mathcal{Z}}^I h^{(1)}]|^2 + |\mathcal{L}_{\mathcal{Z}}^I \mathcal{F}|_{\mathcal{L}\mathcal{N}}^2 + |\mathcal{L}_{\mathcal{Z}}^I \mathcal{F}|_{\mathcal{T}\mathcal{T}}^2 \Big\} w'(q) \, d^3x \, d\tau \\ &+ \varepsilon \sum_{|J'| \leq k-2} \int_0^t \int_{\Sigma_{\tau}} (1+\tau + |q|)^{-1+C\varepsilon} |\nabla [\nabla_{\mathcal{Z}}^{J'} h^{(1)}]|^2 w(q) d^3x \, d\tau \\ &+ \varepsilon \sum_{|J| \leq k-2} \int_0^t \int_{\Sigma_{\tau}} (1+\tau + |q|)^{-1+C\varepsilon} |\mathcal{L}_{\mathcal{Z}}^J \mathcal{F}|^2 w(q) d^3x \, d\tau \\ &+ \varepsilon \sum_{|J| \leq k-1} \int_0^t \int_{\Sigma_{\tau}} (1+\tau + |q|)^{-1+C\varepsilon} |\mathcal{L}_{\mathcal{Z}}^J \mathcal{F}|^2 w(q) d^3x \, d\tau + \varepsilon^3 \end{split}$$

## **Future directions**

- The structure of regularly hyperbolic equations (with Willie Wong)
- Global stability of solutions to the Euler-Einstein system
- Formation of singularities in nonlinear electromagnetic equations

## Thank you