General Relativity and Geometry: Interactions and Missed Opportunities

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Institut Henri Poincaré, Paris November 16, 2015

- All along history, and especially in the 19th and 20th century,
 Geometry and Physics have interacted in a very positive way.
- Taking advantage of the historical dimension of this conference, this lecture presents such instances in the context of the Theory of General Relativity, some have been truly influential, others of the type "missed opportunities".
- Actually, the extent to which this has happened is truly remarkable, and has, in my opinion, not been stressed enough
- As several lectures at this conference will be dedicated to the latest mathematical developments, some very striking and representing remarkable technical achievements, I will concentrate more on the variety of the modes of interaction, stressing the differences in the points of view of physicists and mathematicians that they reveal from an epistemological point of view.

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- Mermann WEYL and Conformal Geometry
- Maluza-Klein Theory, an Anticipation of Bundle Geometry
- Ocrnelius LANCZOS and Generalized Lagrangians
- The ADM Approach to the Einstein Equations
- Supergravity and Killing Spinors
- Some Concluding Remarks

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The Prehistory of General Relativity

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- In spite of the fact that it is dealing with Geometry, it does not contain any figure, and only one formula;
- In the last part entitled "Application on space", B. Riemann discusses the relevance of his considerations to several parts of Physics: Astronomy, the "very large", but also has a paragraph on the "very small";
- He discusses explicitly the relevance of the concepts he introduces to discuss both continuous and discrete settings;
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In a note entitled "On the Space-Theory of Matter" published in 1876, Wiliam Kingdon CLIFFORD develops a truly visionary anticipation of General Relativity. Here is what he says:

- 1) That small portions of space are in fact analogous to little hills on a surface which is on the average flat, namely that the ordinary laws of geometry are not valid in them.
- 2) That this property of being curved or distorted is continually being passed on from one portion of space to another after the manner of a wave.
- 3) That this variation of the curvature of space is what really happens in that phenomenon which we call the motion of matter, whether ponderable or ethereal.
- 4) That in this physical world nothing else takes place but this variation subject to the law of continuity."

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- H. POINCARE's approach, presented in an article entitled "Le mouvement de l'électron" uses invariant theory to identify the geometric structure relevant, namely an indefinite metric of signature (1,3);
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2. Gregorio RICCI-CURBASTRO, a Central Figure for the Development of Geometry

- First in "Delle derivazioni covarianti e controvarianti e del loro uso nella analisi applicati, published in "Studi editi dall'Università di Padova a commemorare l'ottavo centenario della origine della Università di Bologna" in 1888.
- He developed it in the 1892 issue of the Bulletin des Sciences Mathématiques;
- Later, his joint article with his student Tullio LEVI-CIVITA
 "Méthodes de calcul différentiel absolu et leurs applications"
 published in 1900 in the Mathematische Annalen became a
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MÉLANGES.

RÉSUMÉ DE QUELQUES TRAVAUX SUR LES SYSTÈMES VARIABLES DE FONC-TIONS ASSOCIÉS A UNE FORME DIFFÉRENTIELLE QUADRATIQUE (°);

> PAR M. G. RICGI, Professeur à la Faculté des Sciences de Padoue.

4. Notions et définitions générales sur les systèmes variables. — Dans l'Analyse purc aussi bien que dans ses applications à la Géométrie, à la Mécanique et à la Physique, il y a souvent lieu de considérer des systèmes de fonctions de n variables x_1, x_2, \ldots, x_n , qui peuvent toutes être représentées par un symbole général $X_{rs...t}$, dans lequel on a un certain nombre d'indices, qui peuvent tous prendre les n valeurs n, n. Si le nombre des indices est n, nous aurons là ce que nous appellerons un système de fonctions à n variables m^{nple} ou d'ordre n et les n^m fonctions, distinctes ou non, dont il résulte, seront les éléments du système.

2. Les systèmes variables associés à une forme différentielle quadratique. Dérivations covariante et contrevariante selon cette forme.

Dans ce qui va suivre nous allons associer les systèmes variables de fonctions de n variables indépendantes à une expression homogène et du deuxième degré par rapport aux différentielles de ces variables. Nous représenterons par

$$\varphi = \Sigma_{rs} a_{rs} \, dx_r \, dx_s$$

cette expression, qu'on pourra appeler une forme différentielle quadratique à n variables, et nous supposerons qu'elle soit irréductible, c'est-à-dire qu'il n'existe pas une forme différentielle quadratique à m variables y_1, y_2, \ldots, y_m , m étant moindre que n, qui devienne identique à ϖ , en substituant à y_1, y_2, \ldots, y_m certaines fonctions de x_1, x_2, \ldots, x_n . Comme je l'ai démontré ('), pour que la forme φ soit irréductible, il est nécessaire et il suffit que son discriminant, que nous indiquerons par a, ne soit pas nul.

In his 1854 essay B. RIEMANN attaches to a metric g what is now called the *Riemann curvature tensor* R^g , that is a 4-tensor measuring the deviation of a space from being Euclidean.

In "Direzioni et invarianti principali in una varieta qualunque"
 (Atti del Real Inst. Veneto) published in 1904,
 G. RICCI-CURBASTRO introduces the Ricci curvature rg, that he defines as

$$r^{g}(X,Y) = \operatorname{Trace}(Z \mapsto R_{Z,X}^{g}Y)$$
.

- His motivation is purely geometric, namely that of introducing some privileged directions at every point of a Riemannian manifold.
- This motivation proved to be totally useless... but the Ricci curvature proved itself to be a very important concept for completely other reasons.

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5. The Giant Leap Forward of General Relativity

- The shift is considerable as it consists in replacing the scalar field that generated the gravitational force in Newton's theory by a 2-tensor field, measuring generalized lengths, a very considerable mathematical sophistication;
- This requires passing from Minkowski Geometry (the analog of Euclidean Geometry in a purely algebraic setting and the framework of Special Relativity) to Lorentzian Geometry, the analog of Riemannian Geometry.
- If the field equations proposed in their back-to-back articles published in 1913 in the Annalen der Physik involve the Ricci curvature, it does it in a non covariant way, when the theory needs to be fully covariant.

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Consistent field equations were found in 1915 as the result of a joint effort by A. EINSTEIN and David HILBERT.

 As you all know, they obtain these equations as Euler-Lagrange equations of a variational principle for the gravitational potential

$$g \mapsto \int_{\text{space-time}} s^g v_g$$

where $s^g = \text{Trace}_g r^g$ is the scalar curvature of g.

• The Einstein equations are

$$r^g - \frac{1}{2} s^g g = T ,$$

where T is the stress-energy tensor (in the vacuum $T \equiv 0$)

 Several approaches to discuss the Einstein equations have been used since their inception; we will discuss some of ther

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- This was actually a first serious consideration of a gauge theory, and led to further understanding of the general concept of a covariant derivative, not necessarily a metric one
 - By the very way he introduced the scalar field, he was in fact considering conformal classes of metrics, i.e. metrics $\tilde{g} = e^u g$, where u is a smooth function and g a Riemannian or Lorentzian metric.
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5. Kaluza-Klein Theory, an Anticipation of Bundle Geometry

- Starting from a Lorentzian metric g on the space-time M and the electromagnetic potential ξ on M, he introduced an extended 5-dimensional Lorentzian space-time $\tilde{M}=M\times\mathbb{R}$ with metric \tilde{g} obtained using the Ansatz: if $\pi:M\times\mathbb{R}\to M$, then $\tilde{g}=\phi(\pi^*g+(du+\pi^*\xi)\otimes(du+\pi^*\xi))$ (u extra-variable).
- The key formula is the one that describes the 5-dimensional Ricci tensor $r^{\tilde{g}}$ in terms of the data g, ϕ and ξ .
- The striking fact is that claiming that $r^g = 0$ gives back the Maxwell equations for ξ and the right coupled Einstein-Maxwell equations for g et ξ .
- What lies behind is the formalism of *Riemannian submersions* and the remarkable formulas for computing the curvature of the bundle metric from other geometric elements, developed 40 years later by Barrett O'NEILL.

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The Kaluza Ansatz

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It was known since the middle of the 19th century that in dimension 2 the Total Scalar Curvature functional on the space $\mathcal{R}_{iem} M$ of Riemannian metrics on M is constant. Hence:

 It is therefore a differential invariant (actually a topological one) giving rise to the GAUSS-BONNET Formula:

$$\int_{M} s^{g} v_{g} = 2\pi \chi(M) ,$$

where $\chi(M)$ denotes the *Euler-Poincaré characteristic* of M;

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In the late 30s, while looking for other Lagrangians for General Relativity, Cornelius LANCZOS got interested in quadratic functionals in the curvature of the following type on space-times of dimension 4: n

$$\int_{M} (\alpha |R^{g}|^{2} + \beta |r^{g}|^{2} + \gamma (s^{g})^{2}) v_{g}.$$

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7. The Arnowitt-Deser-Misner Approach to the Einstein Equations

- In the ADM approach, first suggested by Yvonne BRUHAT, the 4-dimensional Lorentzian metric g on M is translated into that of a curve $t\mapsto (g_t,k_t,\phi_t)$ where g_t is the induced metric on a space hypersurface Σ , hence a 3-dimensional Riemannian manifold, k_t its second fundamental form and ϕ_t the lapse function.
- The Einstein equations take the form of evolution equations

$$\begin{cases} \frac{\partial g_t}{\partial t} &= -2 \,\phi_t \,k_t \\ \frac{\partial k_t}{\partial t} &= -D^{g_t} d\phi_t + \phi_t \left(r_{g_t} + \left(c_{g_t}(k_t) k_t - 2 \,k_{t \cdot g_t} k_t \right) \right) \end{cases}$$

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and constraint equations

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- This suggests to consider as a primary object the space of Riemannian metrics on a manifold $\mathcal{R}iem\Sigma$, that relativists called for some time Superspace.
- Actually, the Total Scalar Curvature is a function on the space
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- The Einstein equations, the Euler-Lagrange equations of this Lagrangian, express the vanishing of its differential.
- There are usually expressed as the vanishing of the gradient of the function using a metric on Lor M, a space which has several natural Riemannian or pseudo-Riemannian metrics, one of them introduced by Bryce DE WITT.
- Using this approach, the fundamental conservation law $\delta r^g = -\frac{1}{2} ds^g$ expresses that $r^g \frac{1}{2} s^g g$ lies perpendicularly to the orbit of g under diffeomorphims of M, an obvious fact as the Total Scalar Curvature function if invariant under the action of diffeomorphisms.

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- Using this approach, the fundamental conservation law $\delta r^g = -\frac{1}{2} ds^g$ expresses that $r^g \frac{1}{2} s^g$ g lies perpendicularly to the orbit of g under diffeomorphims of M, an obvious fact as the Total Scalar Curvature function if invariant under the action of diffeomorphisms.

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- Hence, this suggests to consider differential equations involving the Ricci curvature in $\mathcal{R}iem\Sigma$, here the second time derivative of the curve of metrics.
- Considering the first derivative is a priori simpler. This is what Thierry AUBIN did, in 1970: he deformed a metric in the direction of its Ricci curvature to improve the curvature.
- The key formula is $\frac{ds^g-u^g}{dt} = |r^g|^2 \frac{1}{2}\Delta_g s^g$, showing for example that, at a constant scalar curvature metric g, this deformation increases the value of the scalar curvature.
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Asking about the Ricci Flow

1979 Berlin Conference on Global Analysis

RICCI CURVATURE and EINSTEIN METRICS

by

Jean Pierre BOURGUIGNON

Centre de Mathématiques Ecole Polytechnique
91128 PALAISEAU CEDEX (France)

- 0.0 After some historical comments, this survey talk intends to summarize what is known about the Ricci curvature of a Riemannian manifold and why it is today especially interesting to consider problems connected with it. Emphasis is put on the relation between the study of the Ricci curvature and the study of Einstein metrics. Since presently we are far from a complete understanding of the situation, we state some problems which, we think, could serve as useful steps towards a better knowledge of this important geometric quantity.
- 3.24 QUESTION -- Does the local flow theorem hold for the vector fields $g \mapsto \mathrm{Ric}^g k \ \mathrm{Scal}^g \ g$ on the space of metrics? What is the global behaviour of their integral curves if they exist?



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- This was done by Richard HAMILTON and Dennis DETURCK in the early 1980s.
- To show that global solutions exist requires that one makes geometric assumptions, and this was done in a series of remarkable papers by R. HAMILTON.
- The rest is History: this has been pushed much further by Grisha PERELMAN, who showed how to go beyond singularities, opening the way to solving the Poincaré Conjecture, a purely topological question.
- Many other applications of the Ricci flow have been made.



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8. Supergravity and Killing Spinors

- Among generalizations of General Relativity, Supergravity stands out as it reappeared a few times after being discarded.
- It combined *Supersymmetry* with General Relativity and fostered the study of very specific dimensions, mostly 7 and 11, with rich geometries such as exceptional *holonomy* G_2 .
- A precise definition of an infinitesimal supersymmetry emerged from these considerations in the form of *Killing Spinors*, i.e. spinor fields ψ satisfying, for any tangent vector X, the equation

$$\nabla_X \psi + \lambda X.\psi = 0,$$

where . denotes Clifford multiplication, and λ is a scalar. The concept was introduced by Martin WALKER and Roger PENROSE studying geodesics of the Kerr metric.

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The purpose of the lecture was to highlight several instances where the Theory of General Relativity

- either triggered the study of new objects, which turned to lead to important developments in Mathematics (bundle geometry, Killing spinors),
- or changed the point of view on others (Ricci curvature),
- or else suggested new approaches (the Ricci flow),
- or even led to missed opportunities

More deeply, as a system of non-linear PDEs with an analytical degeneracy due to its natural geometric nature, the Einstein equations of General Relativity have posed tremendous challenges as far as getting local, and later global solutions, and also for the study of its singularities.

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I thank you for your attention.

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